

Photon/Hadron Discrimination in Hybrid Events at the Pierre Auger Observatory

Philip Ruehl

Universität Siegen

Schule für Astroteilchenphysik 2016, Obertrubach-Bärnfels

06.10.2016

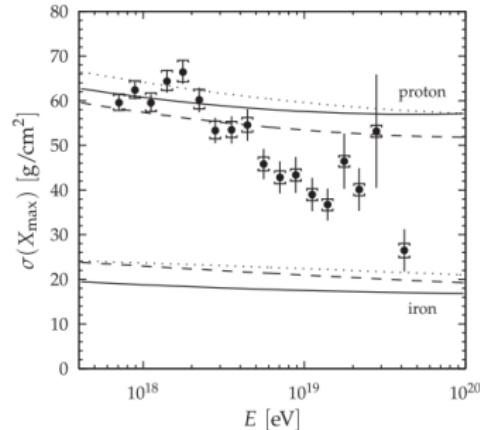
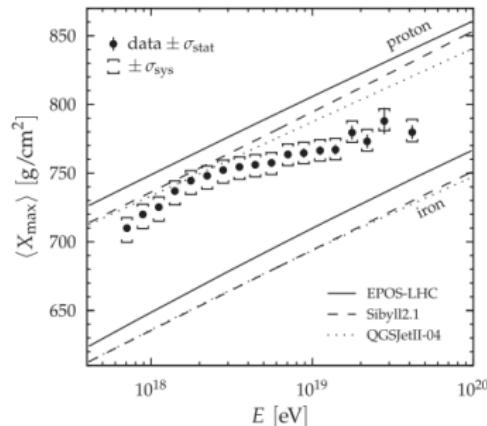


Bundesministerium
für Bildung
und Forschung



Short Digression: Chemical Composition at UHE

[Phys. Rev. D 90 122005]



- UHECR: nuclei from p to Fe
- change of slope suggests heavier particles at higher energies
- X_{max} (later more...) data indicates increase of A with Energy

Why searching for photons...

Models predicting the existence of UHE photons

Top-Down models:

- Z-burst scenario
- super heavy dark matter (SHDM)
- topological defects

Bottom-Up models:

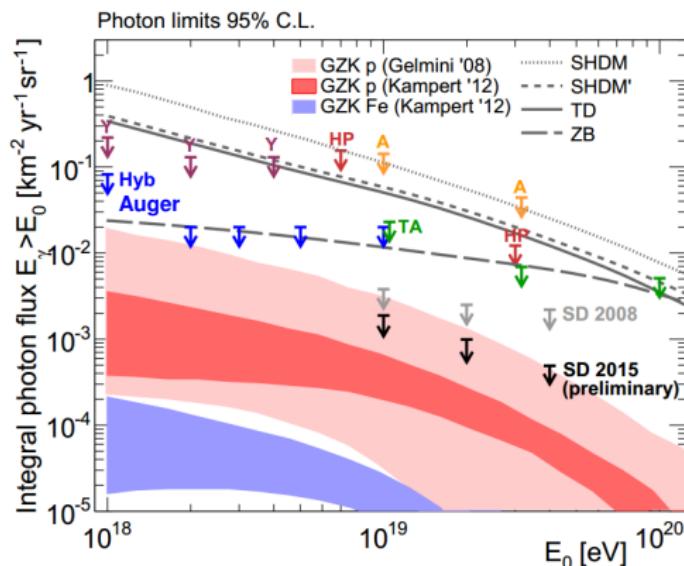
Acceleration of charged nuclear primaries in

- neutron stars / pulsars
- active galactic nuclei
- propagating shock fronts
- gamma ray bursts, etc.

and interaction with CMB ($p_{GZK} + \gamma_{CMB} \rightarrow \Delta^+ \rightarrow \pi^{0(+)} + p(n)$) and $\pi^0 \rightarrow 2\gamma$

Limits on Photon Flux

[C. Bleve for the Auger Collab., ICRC 2015]



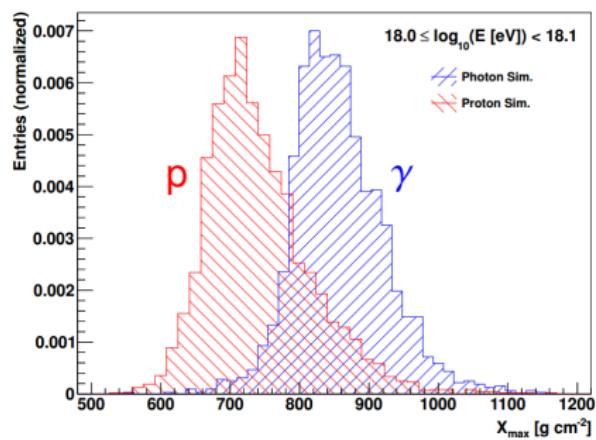
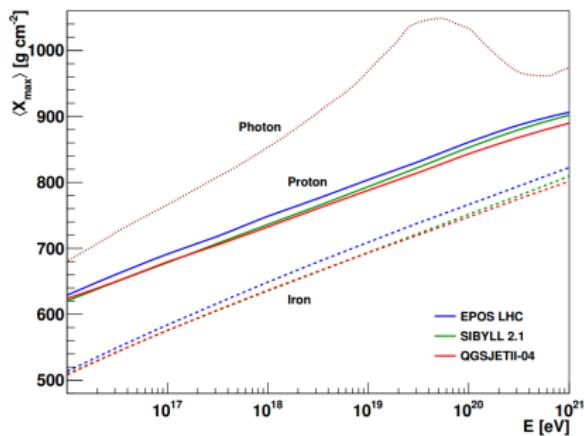
- many Top-Down models severely constrained
- if \exists UHE protons \Rightarrow also bottom-up models predict photon flux caused by the GZK-effect at up to $\sim 1\%$ CR fraction

⇒ Update the limits using hybrid events

Discrimination between UHE photons and nuclei

X_{max} :

- atmospheric depth of the longitudinal shower maximum
- FD observable

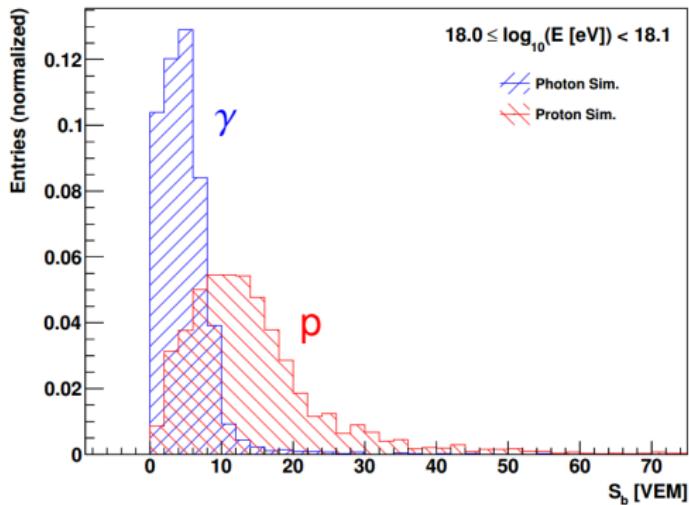


Add SD observables to improve separation power!

SD Observable: S_b

S_b :

- sum of signals S_i in stations with distances r_i from the core
- $S_b := \sum_{i=1}^N \left[S_i \times \left(\frac{r_i}{r_0} \right)^b \right]$
[M. Settimi for the Auger Collab., ICRC 2011]
- uses the difference in the lateral distribution function of showers initiated by different primary types



A New SD Observable F_γ

F_γ uses the **lateral profile** of EAS but takes **advantage of hybrid measurement**

$$F_\gamma := \frac{S_{1000|\gamma}}{S_{1000|Hybrid}}$$

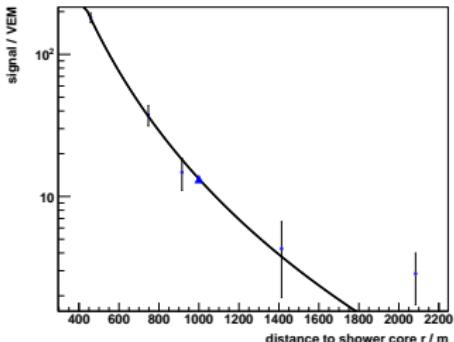
- $S_{1000|\gamma}$: the S_{1000} value obtained from fitting the SD data with a photon-like lateral profile
- $S_{1000|Hybrid}$: the S_{1000} value obtained by using the reconstructed Hybrid energy $E_{FD} \sim S_{1000|Hybrid}^B$ and inverting the standard SD reconstruction.

$$S_{1000|\gamma}$$

Parametrisation of the exponent of the LDF:

$$S(r) = S_{1000|\gamma} \left(\frac{r}{1000m} \right)^{\beta(E,\theta)} \left(\frac{r + 700m}{1700m} \right)^{\beta(E,\theta)}$$

with $\beta(E, \theta)$ from photon simulations (CORSIKA + Offline, QGSJETII-04 + Fluka)



and fitting to data using the maximum likelihood method.

Extended likelihood function:

$$\mathcal{L} = \mathcal{L}_{small} \times \mathcal{L}_{large} \times \mathcal{L}_{non-trig} \times \mathcal{L}_{core} \times \mathcal{L}_{axis} \times \mathcal{L}_{\beta}$$

⇒ only free parameter is $S_{1000|\gamma}$

⇒ in principle applicable to events with only few triggered stations (\rightsquigarrow later more)

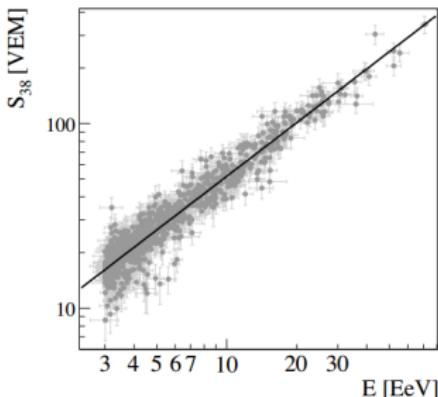
$S_{1000|Hybrid}$

First obtain $S_{38|Hybrid}$ from the reconstructed FD energy

$$S_{38|Hybrid} = \sqrt[B]{\frac{E_{FD}}{A}}$$

with $A = (1.68 \pm 0.05) \times 10^{17} \text{ eV}$ and
 $B = 1.036 \pm 0.009$

[R. Pesce, ICRC 2011]

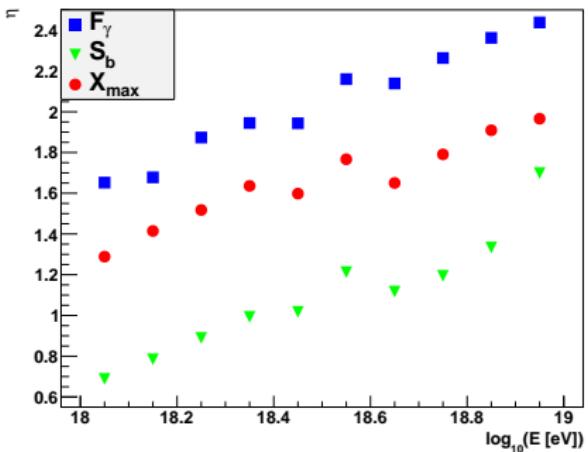
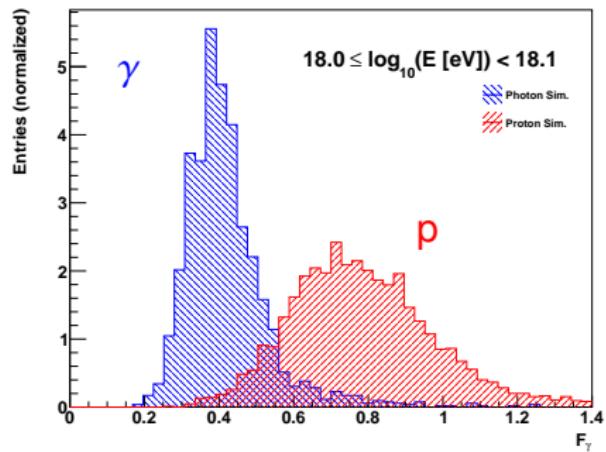


Then obtain $S_{1000|Hybrid}$ by using the reconstructed zenith angle

$$S_{1000|Hybrid} = CIC(\theta) \times S_{38|Hybrid}$$

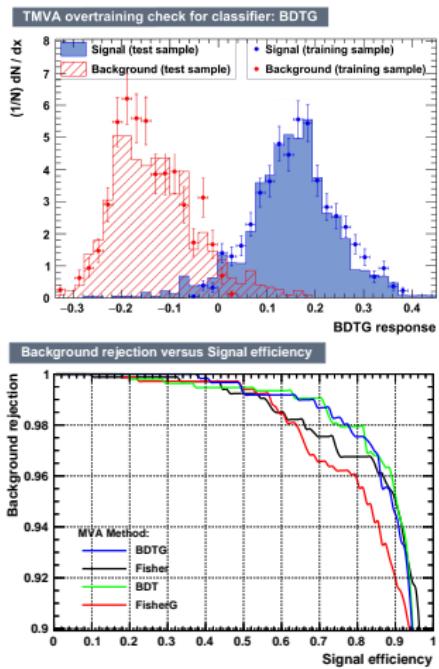
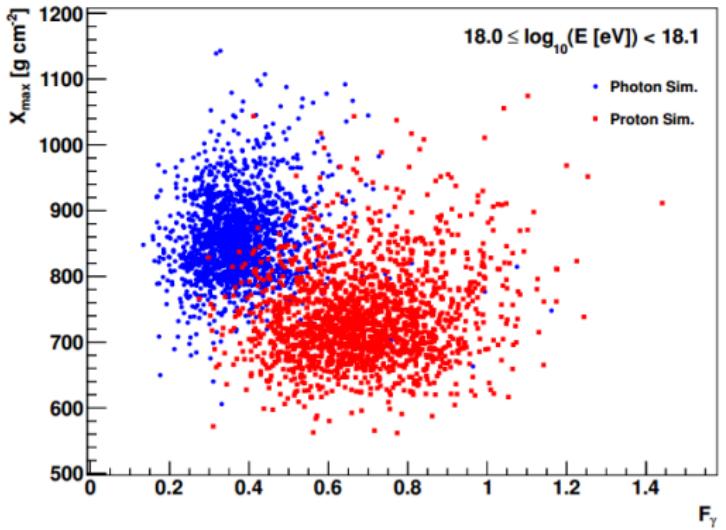
$CIC(\theta)$: 3rd order polynomial in $x = \cos(\theta) - \cos(38^\circ)$ and empiric coefficients

Separation Power of F_γ



~~ high separation power between photon and proton induced air showers!

Combination of F_γ and X_{max} in an MVA



99.5% background reduction at 50% signal efficiency using BDT/BDTG

Summary and outlook

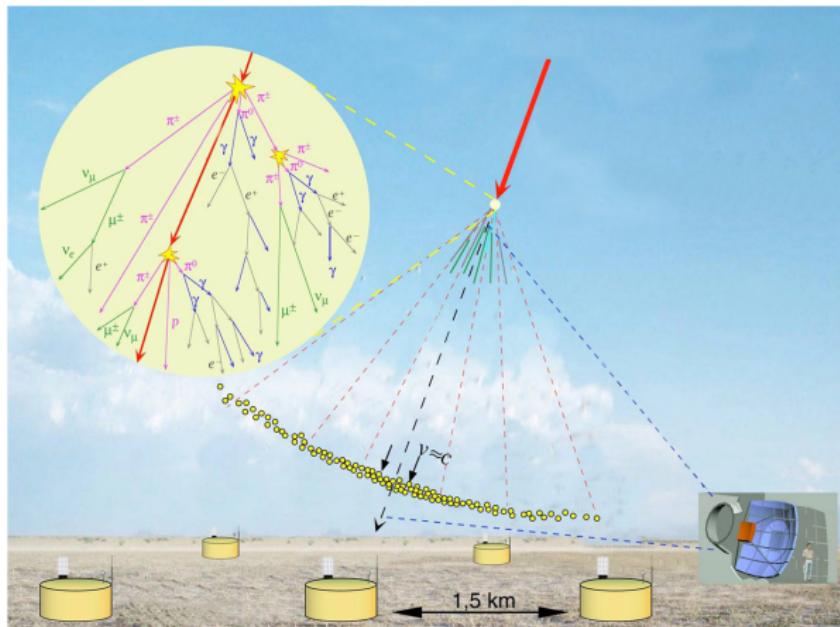
F_γ :

- high separation power between photon and proton induced showers
- useful for photon analysis especially at the lower energies measured at Auger
- combination with X_{max} allows better distinction between primary particles

next step:

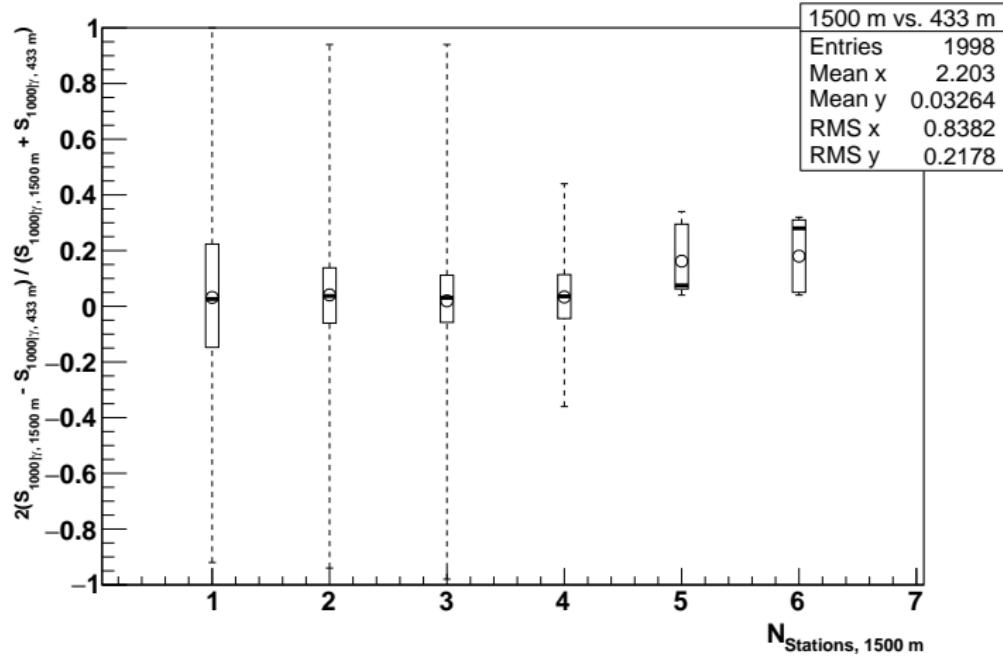
Compare the performances of F_γ and S_b with respect to an MVA to update the upper limits of UHE photon flux starting to probe the GZK region.

Backup - The Pierre Auger Observatory



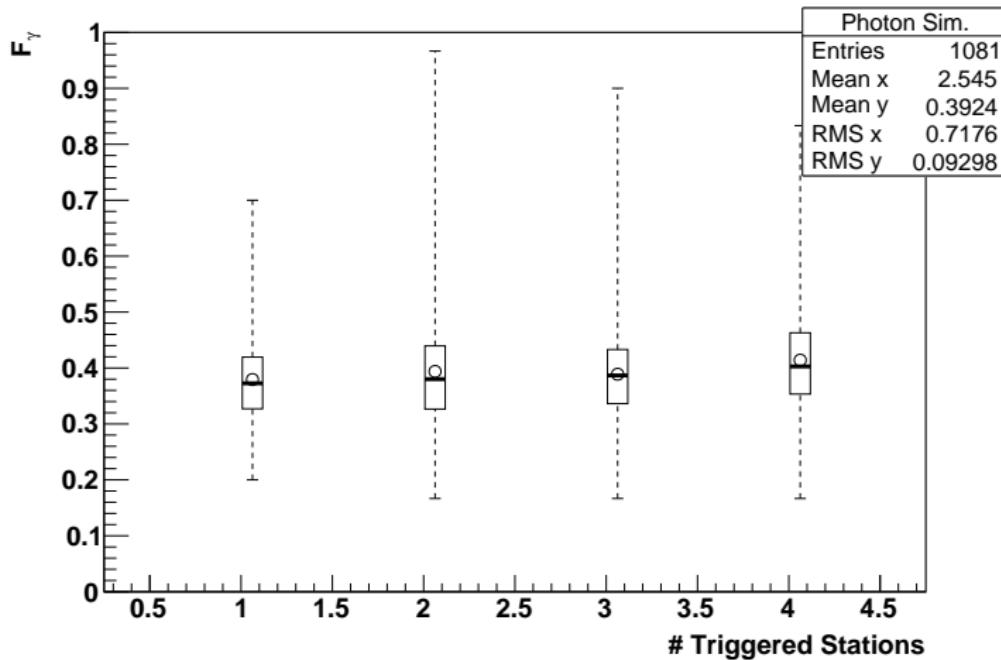
- surface detector (SD): 1660 water cherenkov tanks
- fluorescence detector (FD): 27 fluorescence telescopes at 4 locations
- simultaneous measurements with FD and SD (Hybrid Events)

Comparison With A Dense Array



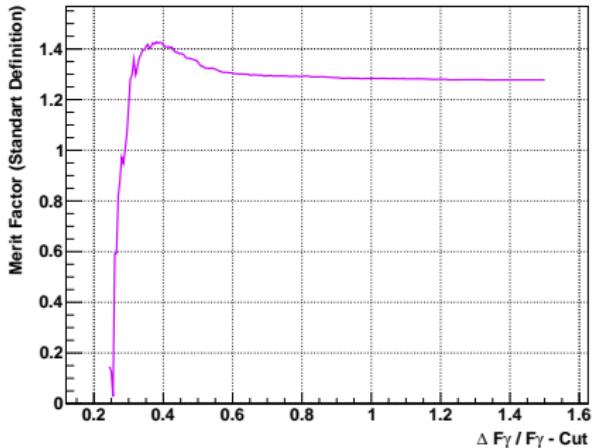
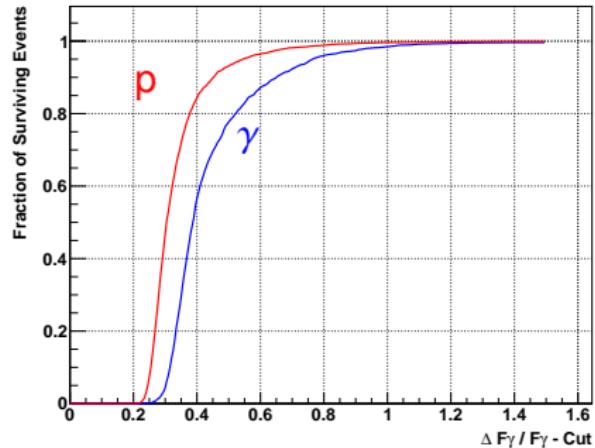
~ Mean value of F_γ is independent of the array geometry

Consistency Check



~~ advantage of F_γ : remains ~constant in events with only one or two triggered stations

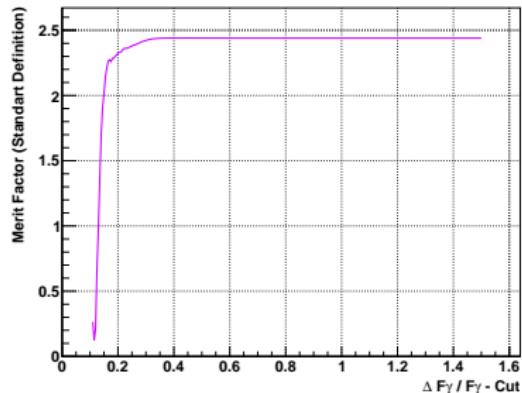
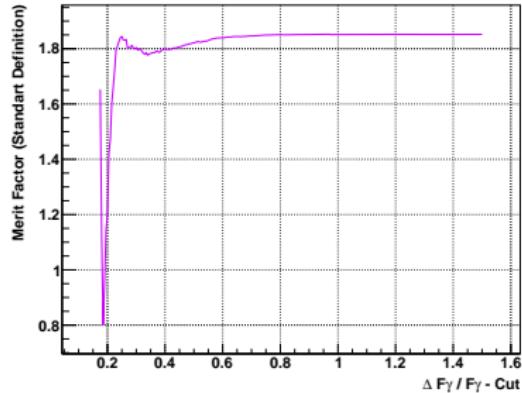
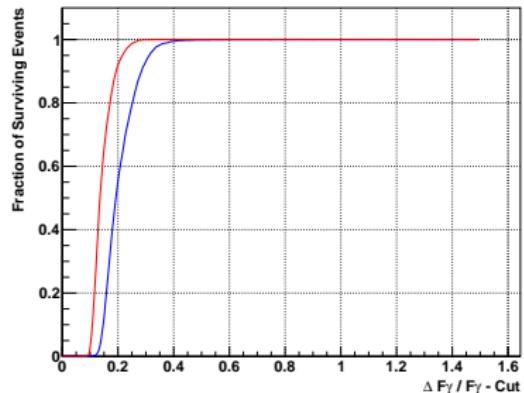
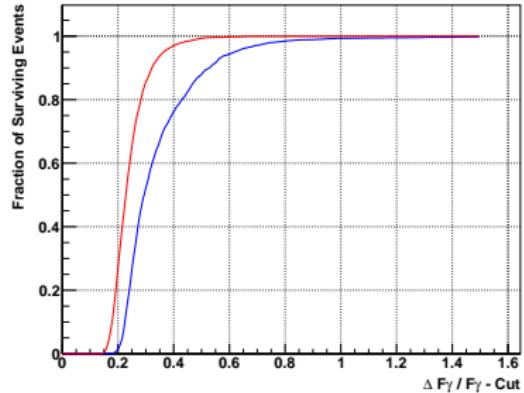
Quality Cut Analysis



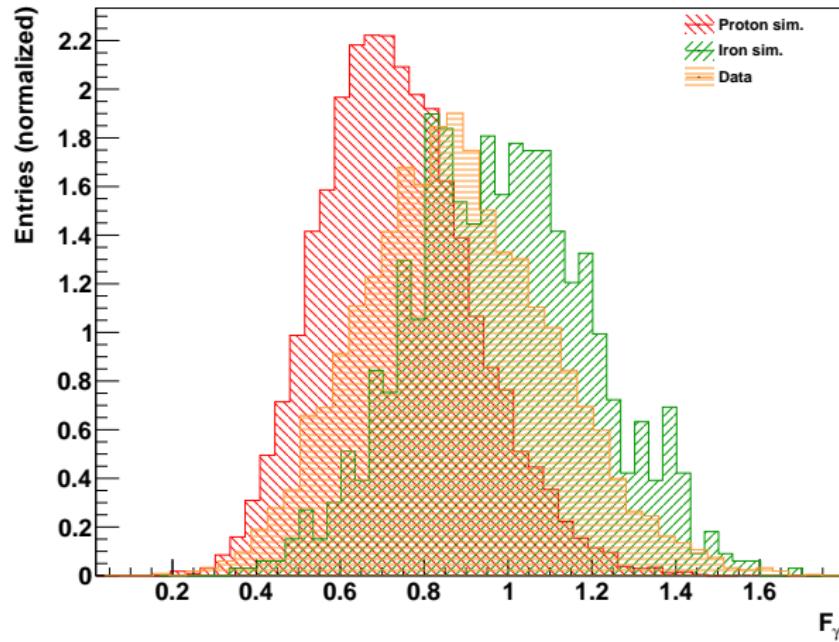
$$\eta := \frac{|\mu_p - \mu_\gamma|}{\sqrt{\sigma_p^2 + \sigma_\gamma^2}}$$

$\Delta F_\gamma / F_\gamma < 0.5 \Rightarrow \text{Efficiency} > 80\%$

Backup - Quality Cut Analysis



Backup - F_γ for heavier nuclei



larger F_γ values for higher nucleon numbers \rightsquigarrow assume pure proton background for conservative estimation.